

Radiation-dominated plasma dynamics in the interaction of super-intense circularly polarized laser pulses with thick plasma targets

Project HRO01. Authors: Tatyana Liseykina^a · Dieter Bauer^a · Andrea Macchi^b · Sergey Popruzhenko^c

Goal, Set-up, Model & Means

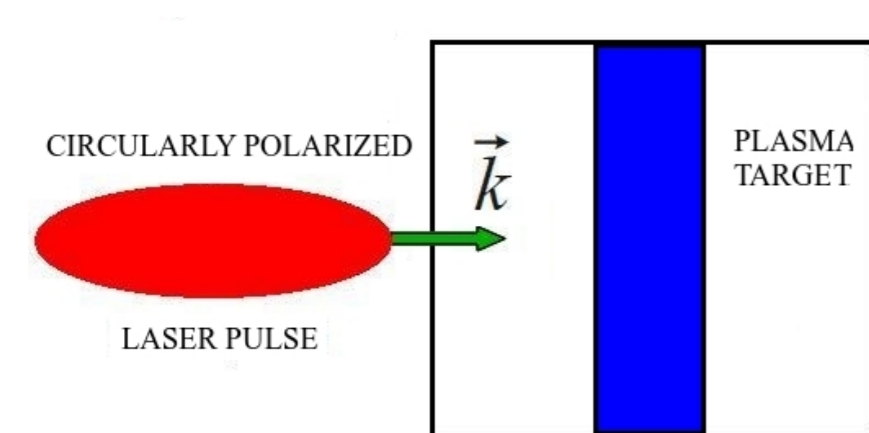
Objectives:

- How strong a magnetic field generated in a laser lab can be?
- Where should we expect the transition between the classical and quantum dynamics in strong-field laser-plasma interactions?

Laser pulse parameters: extremely intense, $I\lambda^2 > 10^{23} \text{ W/cm}^2\mu\text{m}^2$, and circularly polarized

Targets: thickness: much greater than the evanescence length of the laser field; electron density: such to remain above the threshold for relativistic transparency \Rightarrow to keep a steady "hole-boring" regime

Set-up:



Model:

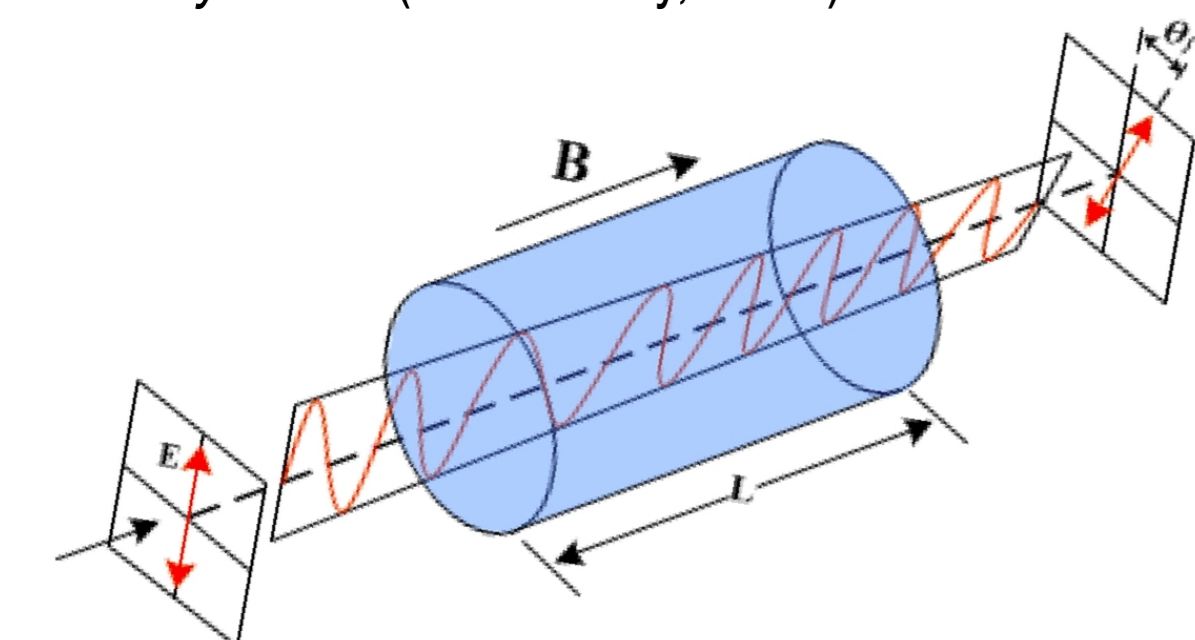
- 3D (essential!)
- Kinetic equation for plasma distribution function
- Maxwell equations for the electromagnetic fields
- Radiation friction included

Means:

- Relativistic Particle-In-Cell simulations

Why circular polarization?

Faraday Effect (M. Faraday, 1845)



Left- and right circularly polarized waves propagate with slightly different phase velocities \rightarrow rotation of the polarization plane of the linear polarized light: $\theta \propto B$.

Inverse Faraday Effect (L. Pitaevskij, 1960)

Circularly polarized light induces a magnetization along the wave vector \vec{k} in the transparent medium. Left- and right-handed polarization waves induce magnetization of opposite signs

- optically transparent plasmas \rightarrow absorption of the angular momentum \rightarrow generation of an axial magnetic field [M.G. Haines, PRL 87 (2001), 135005]

- opaque plasmas???

Logic

- Need to deposit angular momentum into plasma \Rightarrow circular polarization
- Want to achieve extremely strong magnetic fields \Rightarrow extremely strong CP laser pulses
- In intense laser field \Rightarrow radiation is the only mechanism of absorption
 - collisional absorption plays no role
 - collective effects do not lead to absorption of angular momentum
- Should provide a situation when radiation is strong \Rightarrow back reaction of radiation on electrons' dynamics must be taken into account

Hole-boring regime

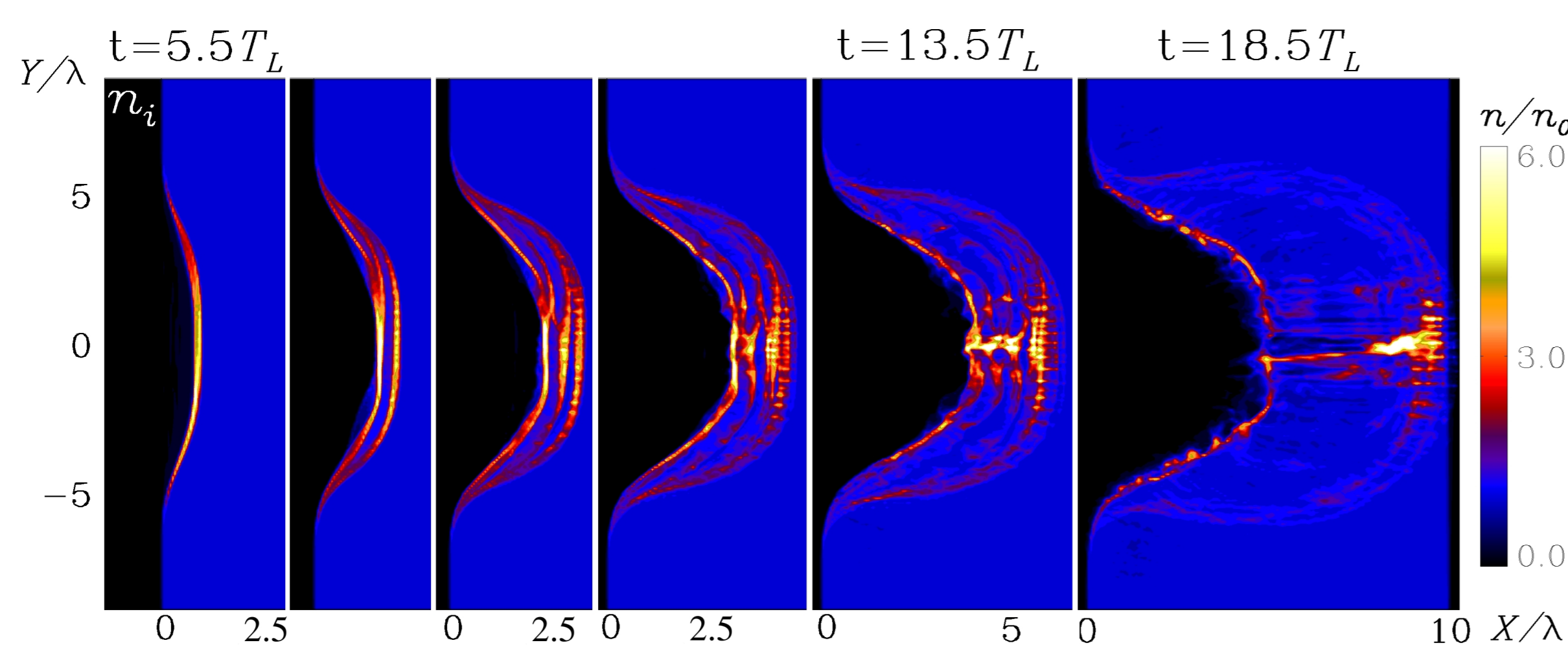


FIG.1. Time evolution of the ion number density for $I_L = 10^{24} \text{ W/cm}^2$, and $n_0 = 1.6 \times 10^{23} \text{ cm}^{-3}$.

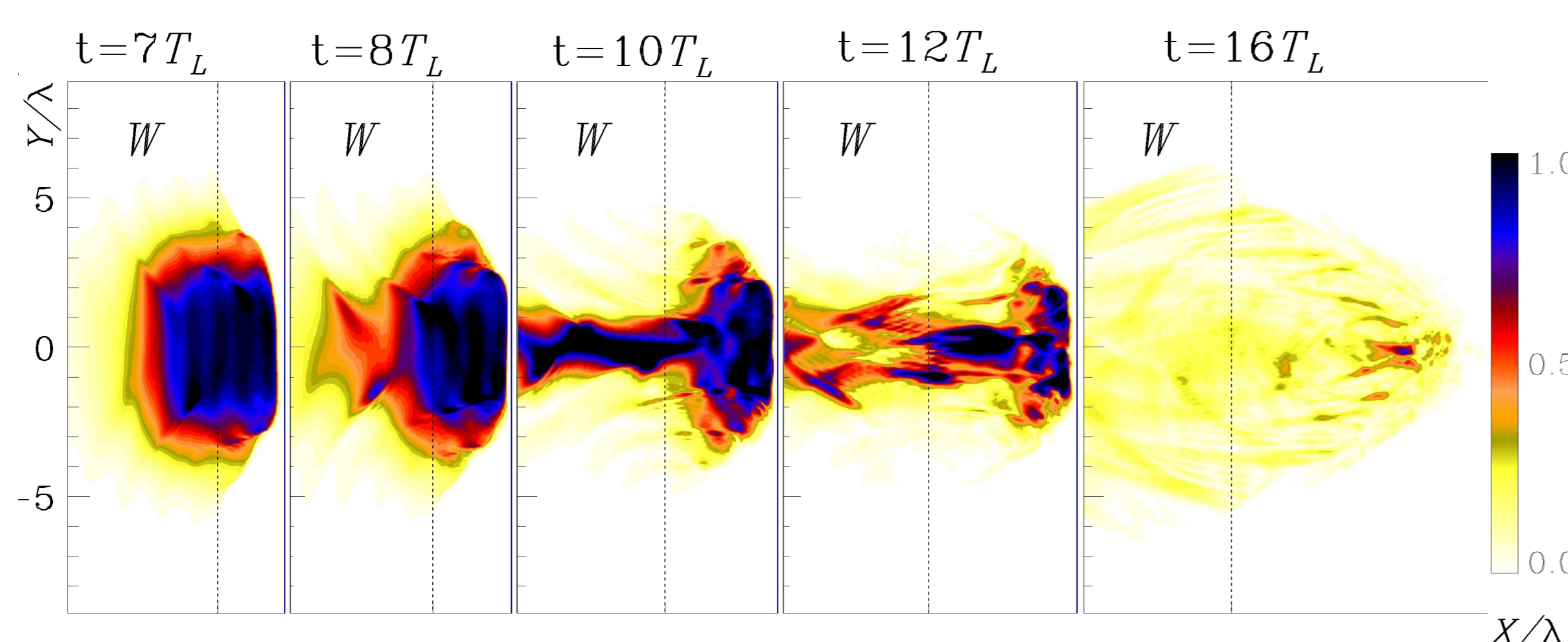


FIG.2. Time evolution of the laser pulser energy density for $I_L = 10^{24} \text{ W/cm}^2$, and $n_0 = 1.6 \times 10^{23} \text{ cm}^{-3}$.

Radiation Reaction modeling

- "Reduced" Landau-Lifshitz equation for electrons

$$\frac{d\mathbf{p}}{dt} = -e \left(\mathbf{E} + \frac{\mathbf{v}}{c} \times \mathbf{B} \right) + \mathbf{F}_{\text{rad}}$$

$$\mathbf{F}_{\text{rad}} = - \left(\frac{2r_e^2}{3} \right) \left\{ \gamma^2 \left[\left(\mathbf{E} + \frac{\mathbf{v}}{c} \times \mathbf{B} \right)^2 - \left(\frac{\mathbf{v}}{c} \cdot \mathbf{E} \right)^2 \right] \frac{\mathbf{v}}{c} - \left[\left(\mathbf{E} + \frac{\mathbf{v}}{c} \times \mathbf{B} \right) \times \mathbf{B} + \left(\frac{\mathbf{v}}{c} \cdot \mathbf{E} \right) \mathbf{E} \right] \right\}$$

more details in M. Tamburini et al., New J. Phys. 10, 123005 (2010)

- Spin force and smaller terms in \mathbf{F}_{rad} are neglected
- Dominant term $\sim \gamma^2 \mathbf{v}$ acts as nonlinear friction

- Assume main contribution to Radiation Friction (RF) losses is at wavelengths $\lambda_{rf} \ll (3/4\pi n_e)^{1/3} \Rightarrow$ radiation is incoherent and escapes from the plasma • it appears as energy dissipation

- RF-relevant wavelengths are not resolved on the numerical grid
- "coherent" wavelengths are double-counted in the RF force, but their contribution is small

Axial Magnetic field with- and without- RF included

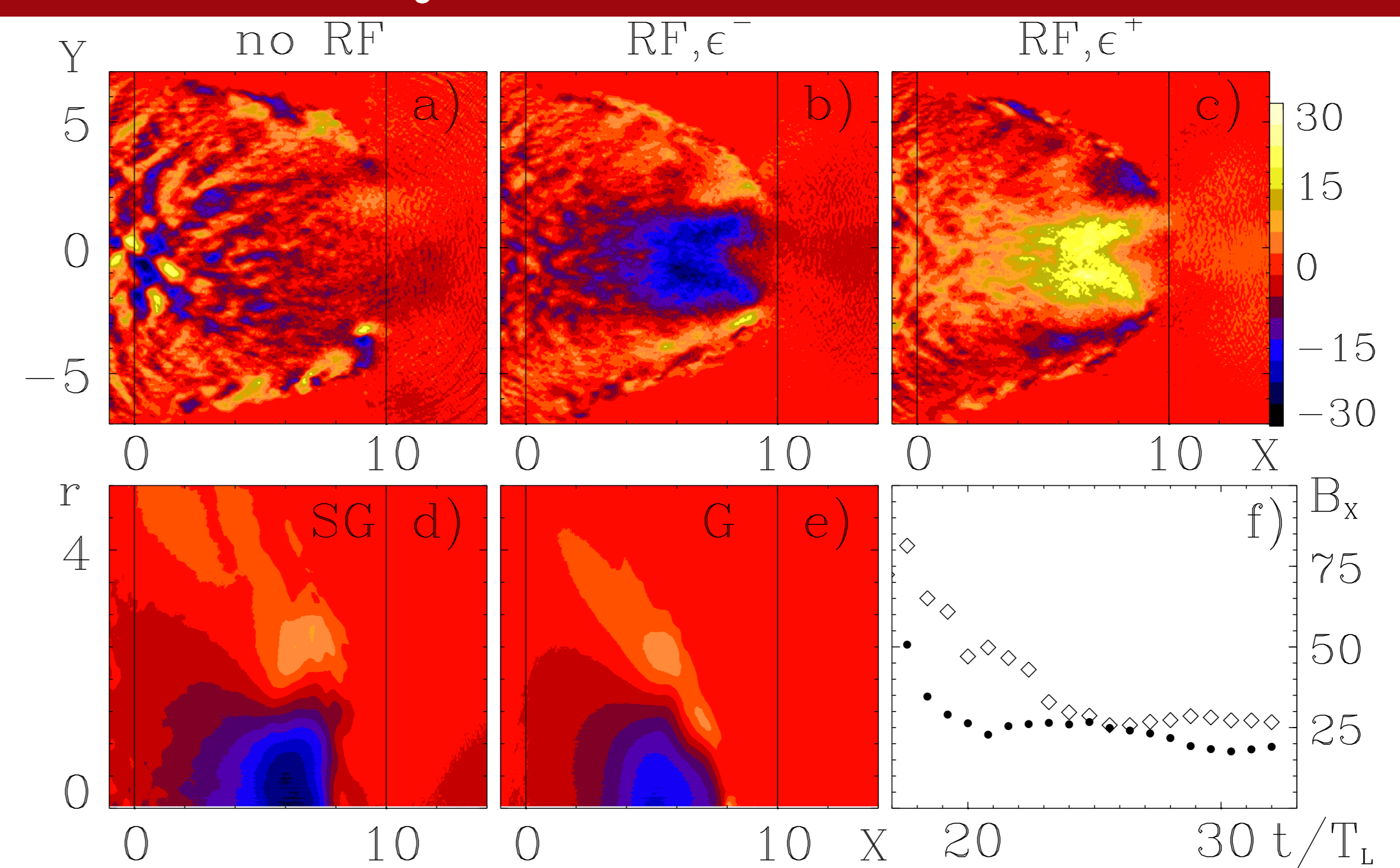
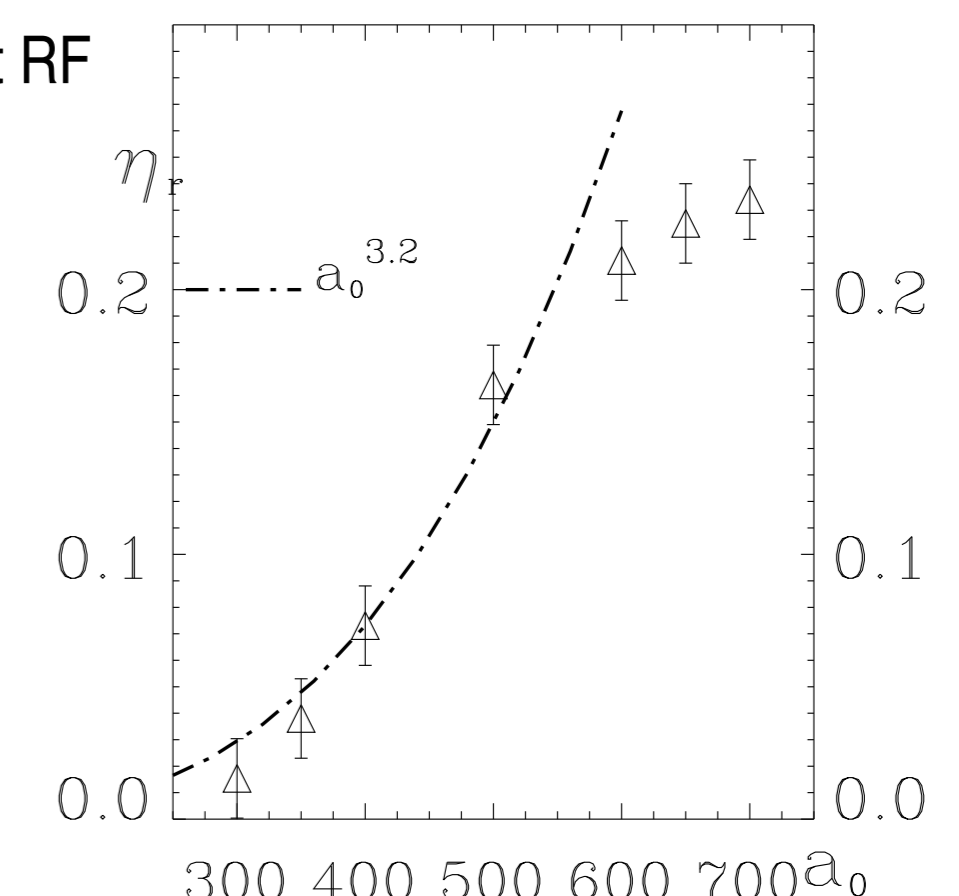


FIG.3. Axial magnetic field B_x normalized to $B_0 = 1.34 \times 10^8 \text{ G}$ a super-Gaussian pulse without RF (a) and with RF included and for opposite helicities (b-c). The field is shown in the xy plane $t = 33\lambda/c$ after the beginning of the interaction. d) and e): B_x averaged over the azimuthal direction comparison for Gaussian (G) and super-Gaussian (SG) pulse profiles. f): the temporal evolution of the maximum value of B_x on the x axis for the G (filled dots) and SG (empty diamonds) pulses.

- no significant axial field is apparent in the simulation without RF
- Axial magnetic field of amplitude $B_{\text{max}} \simeq 4 \text{ GG}$ & extending over several μm is generated
- Sign of the axial magnetic field depends on the polarization
- Field is slowly varying over more than $10 \lambda/c$, no sign of rapid decay

Macroscopic model [1] delivers $B_{\text{max}} \simeq C\eta a_0 B_0$ with η -absorption coefficient:

$$\eta = U_r/U_L = \mathcal{A}c\tau R^2 a_0^2 B_0^2 / \int P_r n_e d^3r dt$$



Radiation in the field of CP wave

$$\vec{E}(t, x) = (E_d, E_0 \cos \varphi, E_0 \sin \varphi), \quad \vec{B}(t, x) = (0, -E_0 \sin \varphi, E_0 \cos \varphi)$$

Model [1]

$$P(\vec{v}) = 2e^2 \omega^2 v_0^2 \gamma^4 / 3c^3 (1 - v_x/c)^2$$

$$P_r \sim a_0^4 \text{ and } U_r \sim a_0^5$$

$$\eta = \xi a_0^3 \text{ with } \xi = 4\pi r_e / 3\lambda \simeq 1.5 \cdot 10^{-8}$$

$$a_{cr} = \xi^{-1/3} \simeq 300 \text{ for } \eta = 1$$

$$I_{cr} \simeq 5 \times 10^{23} \text{ W/cm}^2$$

- Simulations: Correct scaling with intensity but too small absolute values of η

Zeldovich model [2]

$$\vec{F}_r = -P(\vec{v})\vec{v}/c^2$$

$$eE_d - eE_0 (v_0/c) \sin \theta - P(\gamma, v_x) v_x / c^2 = 0$$

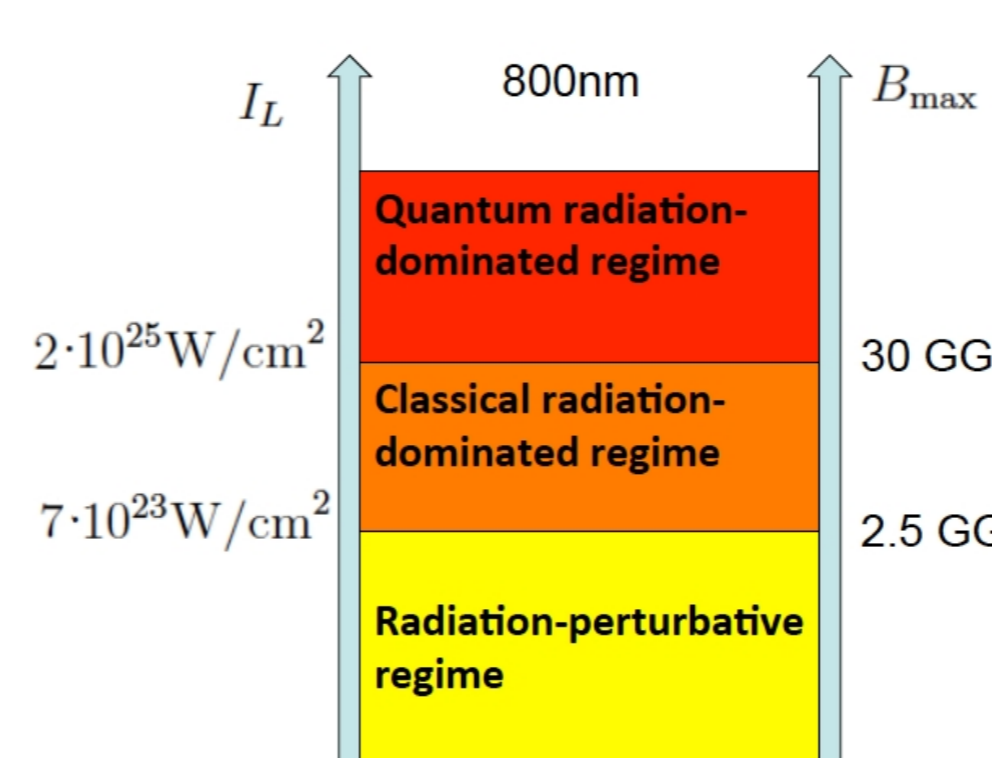
$$eE_0 (1 - v_x/c) \sin \theta + P(\gamma, v_x) v_0 / c^2 = 0$$

$$\gamma m \omega v_0 = eE_0 \cos \theta$$

$$\gamma^2 (1 + \xi^2 \gamma^6) = a_0^2 \text{ and } \eta = \xi \gamma^4 (a_0) / a_0$$

- More consistent with simulations: for $a_0 \gg a_{cr}$, $\eta \simeq 0.5$

Transition from classical to quantum regime [3]



Emitted photon energy \sim electron kinetic energy \rightarrow quantum effects are dominant
 $\hbar\omega_q \simeq \gamma^3 \hbar\omega \simeq \gamma m c^2$
 for $\lambda = 0.8 \mu\text{m} \rightarrow \gamma > \gamma_q \simeq 600$

- Effect of radiation reaction force is not accounted $\rightarrow a_0 < a_q = 600$ and $I < 2 \times 10^{24} \text{ W/cm}^2$
- With $\gamma^2 (1 + \xi^2 \gamma^6) = a_0^2 \rightarrow a_0 < a_q \equiv a(\gamma_q) = 2000$ and $I < 2 \times 10^{25} \text{ W/cm}^2$

1. T. Liseykina, S. Popruzhenko, A. Macchi, New J. Physics 18, 072001 (2016)
2. Y.B. Zeldovich, Usp. Fiz. Nau 115, 161 (1975)
3. S. Popruzhenko, A. Macchi, T. Liseykina, in preparation