

Nuclear Lattice EFT: Status & Perspectives

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- Ulf-G. Meißner, NLEFT: Status and Perspectives - talk, NIC Symposium, Feb. 27, 2020 -

CONTENTS

- Nuclear lattice EFT what and why?
- Chiral EFT on a lattice
- Algorithmic developments
 - the shuttle algorithm
 - the pinhole algorithm
 - the pinhole trace algorithm
- Essentials of nuclear binding
- Ab initio nuclear thermodynamics
- Summary & outlook

Nuclear lattice EFT: what and why?

THE NUCLEAR LANDSCAPE: AIMS & METHODS

- Theoretical methods:
- Lattice QCD: *A* = 0, 1, 2, ...
- NCSM, Faddeev-Yakubowsky, GFMC, ... : A = 3 16
- coupled cluster, . . .: A = 16 100
- density functional theory, ...: $A \ge 10(0)$
- Chiral EFT:
- provides accurate 2N, 3N and 4N forces
- successfully applied in light nuclei with A = 2, 3, 4
 Epelbaum, Hammer, UGM, Rev. Mod. Phys. 81 (2009) 1773
- combine with simulations to get to larger A



\Rightarrow Nuclear Lattice Effective Field Theory

AB INITIO NUCLEAR STRUCTURE and SCATTERING

- Nuclear structure:
 - ★ 3-nucleon forces
 - ★ limits of stability
 - ★ alpha-clustering



- Nuclear scattering: processes relevant for nuclear astrophysics
 - \star alpha-particle scattering: ⁴He + ⁴He \rightarrow ⁴He + ⁴He
 - \star triple-alpha reaction:
 - ★ alpha-capture on carbon:

 4 He + 12 C ightarrow 16 O + γ

 ${}^{4}\text{He} + {}^{4}\text{He} + {}^{4}\text{He} \rightarrow {}^{12}\text{C} + \gamma$

MANY-BODY APPROACHES

- nuclear physics = notoriously difficult problem: strongly interacting fermions
- define *ab initio*: combine the precise and well-founded forces from *chiral EFT* with a many-body approach
- two different approaches followed in the literature:

* combine chiral NN(N) forces with standard many-body techniques

Dean, Duguet, Hagen, Navratil, Nogga, Papenbrock, Schwenk, Soma ...

 \rightarrow successful, but problems with cluster states (SM, NCSM,...)

- * combine chiral forces and lattice simulations methods
- → this new method is called *Nuclear Lattice Effective Field Theory* (NLEFT)

Borasoy, Elhatisari, Epelbaum, Krebs, Lee, Lähde, UGM, Rupak, ...

 \rightarrow rest of the talk

COMPARISON to LATTICE QCD

LQCD (quarks & gluons)	NLEFT (nucleons & pions)	
relativistic fermions	non-relativistic fermions	
renormalizable th'y	EFT	
continuum limit	no continuum limit	
(un)physical masses	physical masses	
Coulomb - difficult	Coulomb - easy	
high T/small $ ho$	small T/nuclear densities	
sign problem severe	sign problem moderate	

- similar methods:
 - hybrid MC, parallel computing, ...
 - \hookrightarrow only treated briefly (shuttle algorithm)
- what I want to discuss within the time limitations:
 - \hookrightarrow how to put the chiral EFT on a lattice
 - \hookrightarrow the pinhole algorithm / center-of-mass in AFQMC
 - \hookrightarrow the pinhole trace algorithm
 - \hookrightarrow going to larger A, increasing the precision
 - \hookrightarrow ab initio nuclear thermodynamics



Chiral EFT on a lattice



T. Lähde & UGM

Nuclear Lattice Effective Field Theory - An Introduction

Springer Lecture Notes in Physics 957 (2019) 1 - 396

NUCLEAR LATTICE EFFECTIVE FIELD THEORY

Frank, Brockmann (1992), Koonin, Müller, Seki, van Kolck (2000), Lee, Schäfer (2004), . . . Borasoy, Krebs, Lee, UGM, Nucl. Phys. **A768** (2006) 179; Borasoy, Epelbaum, Krebs, Lee, UGM, Eur. Phys. J. **A31** (2007) 105

- new method to tackle the nuclear many-body problem
- discretize space-time $V = L_s \times L_s \times L_s \times L_t$: nucleons are point-like particles on the sites
- discretized chiral potential w/ pion exchanges and contact interactions + Coulomb

 \rightarrow see Epelbaum, Hammer, UGM, Rev. Mod. Phys. **81** (2009) 1773

• typical lattice parameters

$$p_{
m max} = rac{\pi}{a} \simeq 314\,{
m MeV}\,[{
m UV}~{
m cutoff}]$$



• strong suppression of sign oscillations due to approximate Wigner SU(4) symmetry

E. Wigner, Phys. Rev. 51 (1937) 106; T. Mehen et al., Phys. Rev. Lett. 83 (1999) 931; J. W. Chen et al., Phys. Rev. Lett. 93 (2004) 242302

ullet physics independent of the lattice spacing for $a=1\dots 2$ fm

Alarcon, Du, Klein, Lähde, Lee, Li, Lu, Luu, UGM, EPJA 53 (2017) 83; Klein, Elhatisari, Lähde, Lee, UGM, EPJA 54 (2018) 121

LATTICE NOTATION

• nucleon annihilation ops:
$$a_{0,0}\equiv a_{\uparrow,p},\ a_{1,0}\equiv a_{\downarrow,p},\ a_{0,1}\equiv a_{\uparrow,n},\ a_{1,1}\equiv a_{\downarrow,n}$$

 \rightarrow labeling **spin** and **isospin**

- spatial & temporal lattice spacing: $a, a_t
 ightarrow lpha_t \equiv a_t/a$
- lattice size: $L \equiv Na, L_t \equiv N_t a_t$

• lattice momenta: $\vec{k} = (k_1, k_2, k_3) \equiv \left(\frac{2\pi}{N}\hat{k}_1, \frac{2\pi}{N}\hat{k}_2, \frac{2\pi}{N}\hat{k}_3\right),$

ightarrow in the first Brillouin zone: $|k_i| < \pi$ and $0 \leq |\hat{k}_i| < N/2$

 any derivative operator requires *improvement*, as the simplest representation in terms of two neighboring points is afflicted by the largest discretization errors

$$\begin{split} k_l &\equiv \sum_{j=1}^{\nu+1} (-1)^{j+1} \theta_{\nu,j} \sin(jk_l) + \mathcal{O}(a^{2\nu+2}) \\ \frac{k_l^2}{2} &\equiv \sum_{j=0}^{\nu+1} (-1)^j \omega_{\nu,j} \cos(jk_l) + \mathcal{O}(a^{2\nu+2}) \end{split}$$

 $\hookrightarrow \text{ no improvement } (\nu=0) \text{:} \ \ \theta_{0,1}=1, \ \ \omega_{0,0}=1, \ \ \omega_{0,1}=1$



LATTICE NOTATION continued

• Order
$$a^2$$
 improvement ($\nu = 1$): $\theta_{1,1} = \frac{4}{3}, \ \theta_{1,2} = \frac{1}{6}, \ \omega_{1,0} = \frac{5}{4}, \ \omega_{1,1} = \frac{4}{3}, \ \omega_{1,2} = \frac{1}{12}$

• Order
$$a^4$$
 improvement ($\nu = 2$): $\theta_{2,1} = \frac{3}{2}, \ \theta_{2,2} = \frac{3}{10}, \ \theta_{2,3} = \frac{1}{30}$
 $\omega_{2,0} = \frac{49}{36}, \ \omega_{2,1} = \frac{3}{2}, \ \omega_{2,2} = \frac{3}{20}, \ \omega_{2,3} = \frac{1}{90}$

 \hookrightarrow definition of the first order spatial derivative:

$$abla_{l,(
u)}f(ec{n}) \equiv rac{1}{2}\sum_{j=1}^{
u+1} (-1)^{j+1} heta_{
u,j} igg[f(ec{n}+j\hat{e}_l)-f(ec{n}-j\hat{e}_l)igg]$$

 \hookrightarrow second order spatial derivative:

$$ilde{
abla}_{l,(
u)}^2 f(ec{n}) \equiv -\sum_{j=0}^{
u+1} (-1)^j \omega_{
u,j} igg[f(ec{n}+j\hat{e}_l) + f(ec{n}-j\hat{e}_l) igg]$$

has two zeros in per Brillouin zone \rightarrow beneficial feature for tuning NLO coefficients

$$\hookrightarrow \text{ improved lattice dispersion relation: } \omega^{(\nu)}(\vec{p}) \equiv \frac{1}{\tilde{m}_N} \sum_{j=0}^{\nu+1} \sum_{l=1}^3 (-1)^j \omega_{\nu,j} \, \cos(jp_l) \\ \\ \tilde{m}_N \equiv m_N a$$

TRANSFER MATRIX METHOD

- Correlation-function for A nucleons: $Z_A(\tau) = \langle \Psi_A | \exp(-\tau H) | \Psi_A \rangle$ with Ψ_A a Slater determinant for A free nucleons [or a more sophisticated (correlated) initial/final state]
- Transient energy

$$E_A(au) = -rac{d}{d au}\,\ln Z_A(au)$$

 \rightarrow ground state: $E_A^0 = \lim_{\tau \to \infty} E_A(\tau)$

• Exp. value of any normal–ordered operator ${\cal O}$

$$Z_A^{\mathcal{O}} = raket{\Psi_A} \exp(- au H/2) \, \mathcal{O} \, \exp(- au H/2) \ket{\Psi_A}$$

$$\lim_{ au o \infty} \, rac{Z_A^{\mathcal{O}}(au)}{Z_A(au)} = \langle \Psi_A | \mathcal{O} \, | \Psi_A
angle \, ,$$

Euclidean time



Euclidean time

CONFIGURATIONS







 $\begin{array}{l} \Rightarrow \text{ all } \textit{possible} \text{ configurations are sampled} \\ \Rightarrow \text{ preparation of } \textit{all possible} \text{ initial/final states} \\ \Rightarrow \textit{clustering} \text{ emerges } \textit{naturally} \end{array}$

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AUXILIARY FIELD METHOD

• Represent interactions by auxiliary fields:



COMPUTATIONAL EQUIPMENT

- Past = JUQUEEN (BlueGene/Q)
- Present = JUWELS (modular system) + SUMMIT + ...





Algorithmic developments

shuttle algorithm: Lu et al., Phys. Lett. B797 (2019) 134863pinhole algorithm: Elhatisari et al., Phys. Rev. Lett. 19 (2017) 222505pinhole trace algorithm: Lu et al., arXiv:1912.05105, subm. for publication

THE SHUTTLE ALGORITHM

- Auxiliary fields $s(n_t, \vec{n})$ are updated on one time slice only
- Proceed to the next time slice, update, ..., turn around at the end of the time series
- \rightarrow very efficient for small temporal lattice spacings, $a_t=0.001~{\rm MeV^{-1}}$
- ightarrow high acceptance rate, typically $\sim 50\%$
- \rightarrow more efficient than HMC
- \rightarrow about 10 times more configurations per hour generated
- \rightarrow further acceleration by GPUs





CENTER-of-MASS PROBLEM

 AFQMC calculations involve states that are superpositions of many different center-of-mass (com) positions

 $egin{aligned} Z_A(au) &= \langle \Psi_A(au) | \Psi_A(au)
angle \ &| \Psi_A(au)
angle &= \exp(-H au/2) | \Psi_A
angle \end{aligned}$



• but: translational invariance requires summation over all transitions

 $Z_A(au) = \sum_{i_{
m com}, j_{
m com}} \langle \Psi_A(au, i_{
m com}) | \Psi_A(au, j_{
m com})
angle, \ \ {
m com} = {
m mod}((i_{
m com} - j_{
m com}), L)$

 $i_{\rm com}~(j_{\rm com})=$ position of the center-of-mass in the final (initial) state

- \rightarrow density distributions of nucleons can not be computed directly, only moments
- \rightarrow need to overcome this deficieny

PINHOLE ALGORITHM

Solution to the CM-problem:

track the individual nucleons using the *pinhole algorithm*

 Insert a screen with pinholes with spin & isospin labels that allows nucleons with corresponding spin & isospin to pass = insertion of the A-body density op.:

$$egin{aligned} &
ho_{i_1,j_1,\cdots i_A,j_A}(\mathrm{n}_1,\cdots \mathrm{n}_A)\ &=:
ho_{i_1,j_1}(\mathrm{n}_1)\cdots
ho_{i_A,j_A}(\mathrm{n}_A): \end{aligned}$$

MC sampling of the amplitude:

$$\begin{array}{l} \text{MC sampling of the amplitude:} & & & \\ A_{i_1,j_1,\cdots i_A,j_A}(\mathbf{n}_1,\ldots,\mathbf{n}_A,L_t) & & \\ = \langle \Psi_A(\tau/2) | \rho_{i_1,j_1,\cdots i_A,j_A}(\mathbf{n}_1,\ldots,\mathbf{n}_A) | \Psi_A(\tau/2) \rangle \end{array}$$

- Allows to measure proton and neutron distributions
- Resolution scale $\sim a/A$ as cm position $\mathbf{r_{cm}}$ is an integer $\mathbf{n_{cm}}$ times a/A

 $\tau_i = \tau$

 $\tau/2$

 $\tau_i =$



PROTON and NEUTRON DENSITIES in CARBON

- first NLEFT calculation of the charge density in ¹²C [proton size accounted for]
- asymptotic properties of the distributions from the volume dependence of N-body bound states
 König,

König, Lee, Phys. Lett. B779 (2018) 9

• open symbols: neutron / closed symbols: proton



 \Rightarrow independent of projection time \rightarrow ground state \Rightarrow small error bars \rightarrow sign problem under control

FORM FACTORS

- Fit charge distributions by a Wood-Saxon shape
 - \hookrightarrow get the form factor from the Fourier-transform (FT)
 - \hookrightarrow uncertainties from a direct FT of the lattice data



 \Rightarrow detailed structure studies become possible

PINHOLE TRACE ALGORITHM (PTA)



$$=\sum_{n_1,\cdots,n_A}\int \mathcal{D}s\mathcal{D}\pi\langle n_1,\cdots,n_A|\exp[-\beta H(s,\pi)]|n_1,\cdots,n_A\rangle$$

 allows to study: liquid-gas phase transition → this talk thermodynamics of finite nuclei
 thermal dissociation of hot nuclei
 cluster yields of dissociating nuclei Essentials of Nuclear Binding

B. N. Lu, N. Li, S. Elhatisari, D. Lee, E. Epelbaum, UGM, Phys. Lett. **B 797** (2019) 134863

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TOWARDS HEAVY NUCLEI in NLEFT

- Two step procedure:
 - 1) Further improve the LO action
 - \hookrightarrow minimize the sign oscillations
 - \hookrightarrow minimize the higher-body forces
 - \hookrightarrow gain an understanding of the essentials of nuclear binding

2) Work out the corrections to N3LO

- \hookrightarrow first on the level of the NN interaction
- \hookrightarrow second for the spectra of nuclei
- \hookrightarrow third for nuclear reactions (nuclear astrophysics)

ESSENTIAL ELEMENTS for NUCLEAR BINDING I

Lu, Li, Elhatisari, Epelbaum, Lee, UGM. Phys. Lett. B 797 (2019) 134863 [arXiv:1812.10928]

- Step 1: construct a highly improved LO action, free of any sign problem
- Highly SU(4) symmetric LO action without pions, local and non-local smearing:

$$\begin{split} H_{\mathrm{SU}(4)} &= H_{\mathrm{free}} + \frac{1}{2!} C_2 \sum_n \tilde{\rho}(n)^2 + \frac{1}{3!} C_3 \sum_n \tilde{\rho}(n)^3 \\ \tilde{\rho}(n) &= \sum_i \tilde{a}_i^{\dagger}(n) \tilde{a}_i(n) + \frac{s_L}{|n'-n|=1} \sum_i \tilde{a}_i^{\dagger}(n') \tilde{a}_i(n') \\ \tilde{a}_i(n) &= a_i(n) + \frac{s_{NL}}{|n'-n|=1} \sum_i a_i(n') \\ &+ \frac{1}{|n'-n|=1} \sum_{|n'-n|=1} a_i(n') + \frac{1}{|n'-n|=1} \sum_{|n'-n|=1} a_i(n') \\ &+ \frac{1}{|n'-n|=1} \sum_{|n'-n|=1} a_i(n') + \frac{1}{|n'-n|=1} \sum_{|n'-n|=1} a_i(n') \\ &+ \frac{1}{|n'-n|=1} \sum_{|n'-n|=1} a_i(n') + \frac{1}{|n'-n|=1} \sum_{|n'-n|=1} a_i(n') \\ &+ \frac{1}{|n'-n|=1} \sum_{|n'-n|=1} a_i(n') + \frac{1}{|n'-n|=1} \sum_{|n'-n|=1} a_i(n') \\ &+ \frac{1}{|n'-n|=1} \sum_{|n'-n|=1} a_i(n') + \frac{1}{|n'-n|=1} \sum_{|n'-n|=1} a_i(n') \\ &+ \frac{1}{|n'-n|=1} \sum_{|n'-n|=1} a_i(n') + \frac{1}{|n'-n|=1} \sum_{|n'-n|=1} a_i(n') \\ &+ \frac{1}{|n'-n|=1} \sum_{|n'-n|=1} a_i(n') + \frac{1}{|n'-n|=1} \sum_{|n'-n|=1} a_i(n') \\ &+ \frac{1}{|n'-n|=1} \sum_{|n'-n|=1} a_i(n') + \frac{1}{|n'-n|=1} \sum_{|n'-n|=1} a_i(n') \\ &+ \frac{1}{|n'-n|=1} \sum_{|n'-n|=1} a_i(n') + \frac{1}{|n'-n|=1} \sum_{|n'-n|=1} a_i(n') \\ &+ \frac{1}{|n'-n|=1} \sum_{|n'-n|=1} a_i(n') + \frac{1}{|n'-n|=1} \sum_{|n'-n|=1} a_i(n') \\ &+ \frac{1}{|n'-n|=1} \sum_{|n'-n|=1} a_i(n') + \frac{1}{|n'-n|=1} \sum_{|n'-n|=1} a_i(n') \\ &+ \frac{1}{|n'-n|=1} \sum_{|n'-n|=1} \sum_{|n'-n|=1} a_i(n') \\ &+ \frac{1}{|n'-n|=1} \sum_{|n'-n|=1} \sum_{|n'-n|=1}$$

• Only four parameters!

 C_2 and C_3 = strength of the leading two- and three-body interactions s_L and s_{NL} = strength of the local and the non-local interaction



ESSENTIAL ELEMENTS for NUCLEAR BINDING II

- Fixing the parameters:
 - \star interaction strength C_2 and range s_L from the average S-wave scattering lengths and effective ranges (requires SU(4) breaking later)
 - \star interaction strength C_3 from the ³H binding energy
 - \star interaction range s_{NL} can not be determined in light nuclei
 - \hookrightarrow calculate the volume- and surface energy of mid-mass nuclei $16 \leq A \leq 40$
 - compare w/ existing calculations:



Mac-Mic: Wang et al., Phys. Lett. B **734** (2014) 215 FRLDM: Möller et al., Atom Data Nucl. Data Tabl. **59** (1995) 184 mean field: Bender et al., Rev. Mod. Phys. **75** (2003) 121



ENERGIES for SELECTED NUCLEI

 Calculated binding energies for 3N & alpha-type nuclei: Binding energies for 86 even-even nuclei



- selected nuclei: amazingly precise, largest deviation about 4% in ¹⁶O
- even-even isotopic chains come out amazingly precise, general trends reproduced
- \hookrightarrow on the proton-rich side better than on the neutron-rich one \rightarrow spin-dep. effects
- but remember: this is only leading order!

RADII for SELECTED NUCLEI

 Calculated charge radii for 3N & alpha-type nuclei:

	$R_{ m ch}$	Exp.	$R_{ m ch}/{ m Exp.}$
³ Н	1.90(1)	1.76	1.08
³ He	1.99(1)	1.97	1.01
⁴ He	1.72(3)	1.68	1.02
^{16}O	2.74(1)	2.70	1.01
²⁰ Ne	2.95(1)	3.01	0.98
^{24}Mg	3.13(2)	3.06	1.02
²⁸ Si	3.26(1)	3.12	1.04
⁴⁰ Ca	3.42(3)	3.48	0.98

 Charge distributions for ¹⁶O and ⁴⁰Ca



- Radii quite well described
- ↔ overcomes earlier problems (see PRL 109 (2012) 252501, 112 (2014) 102501)
- Also a fair description of the charge distributions at LO!

NEUTRON MATTER

\bullet 14 to 66 neutrons in $L=5, 6, 7 ightarrow ho=0.02-0.15\,{ m fm^{-3}}$



• exact SU(4)

 \hookrightarrow deviations at low densities

• SU(4) breaking term $\rightarrow a_{nn} \sqrt{}$ \hookrightarrow good overall description

APR = Akmal, Pandharipande, Ravenhall, Phys. Rev. C 58 (1998) 1804; GCR = Gandolfi, Carlson, Reddy, Phys. Rev. C 85 (2012) 032801; all others in: Tews et al., Phys. Rev. Lett. 110 (2013) 032504.

GOING to HIGHER ORDERS

• Step 2: Work out NN phase shifts to N3LO

 \hookrightarrow calculate nuclear properties based on NN forces only \surd

 \hookrightarrow add three-nucleon forces (in the works)

• Starting Hamiltonian:

 $H_{2N} = H_{\text{free}} + H_{SU(4)}(C_3 = 0) + H_{OPE}$

 \hookrightarrow consistent with the power counting (3NFs appear at N2LO)

 \hookrightarrow OPE induces some sign problem (but manageable)

 NN forces to N3LO worked out for 4 latt. spacings a = 0, 99, 1.32, 1.64, 1.98 fm with local and non-local smearing and restoration of Galilean invariance breaking

 \hookrightarrow used as input for nuclear structure calculations

Li et al., Phys. Rev. C 98 (2018) 044002; Phys. Rev. C 99 (2019) 064001

NN INTERACTION at N3LO

NLO

N₂LO

N3LO

• np phase shifts including uncertainties for a = 1.64 fm (cf. Nijmegen PWA)



uncertainty estimates á la Epelbaum, Eur. Phys. J. A **51** (2015) 53

UGM,

Krebs,

NUCLEI at N3LO

\bullet Binding energies of nuclei for $a=1.64\,{\rm fm}$



\rightarrow excellent starting point for precision studies

Ab Initio Nuclear Thermodynamics

B. N. Lu, N. Li, S. Elhatisari, D. Lee, J. Drut, T. Lähde, E. Epelbaum, UGM, [arXiv:1912.05105], submitted for publication

PHASE DIAGRAM

Fig. courtesy B.-N. Lu

• Phase diagram of strongly interacting matter



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NEW PARADIGM for NUCLEAR THERMODYNAMICS

- The PTA allows for simulations with fixed neutron & proton numbers at non-zero T
- \hookrightarrow thousands to millions times faster than existing codes using the grand-canonical ensemble ($t_{
 m CPU} \sim V N^2$ vs. $t_{
 m CPU} \sim V^3 N^2$)
- \bullet Only a mild sign problem \rightarrow pinholes are dynamically driven to form pairs
- Typical simulation parameters:

up to N = 144 nucleons in volumes $L^3 = 4^3, 5^3, 6^3$ \hookrightarrow densities from 0.008 fm⁻³ ... 0.20 fm⁻³ a = 1.32 fm $\rightarrow \Lambda = \pi/a = 470$ MeV , $a_t \simeq 0.1$ fm consider $T = 10 \dots 20$ MeV

 \bullet use twisted bc's, average over twist angles \rightarrow acceleration to the td limit

• very favorable scaling for generating config's:

$$\Delta t \sim N^2 L^3$$

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CHEMICAL POTENTIAL

• Calculated from the free energy: $\mu = (F(N+1) - F(N-1))/2$



⁻ Ulf-G. Meißner, NLEFT: Status and Perspectives - talk, NIC Symposium, Feb. 27, 2020 -

EQUATION of STATE

• Calculated by integrating: $dP = \rho \, d\mu$

• Crtitical point: $T_c = 15.8(1.6)$ MeV, $P_c = 0.26(3)$ MeV/fm³, $\rho_c = 0.089(18)$ fm⁻³



⁻ Ulf-G. Meißner, NLEFT: Status and Perspectives - talk, NIC Symposium, Feb. 27, 2020 -

3

fm

0.06(2)

 $ho_{\rm c}$

0.31(7) MeV/fm³

VAPOR-LIQUID PHASE TRANSITION

- Vapor-liquid phase transition in a finite volume $V \ \& \ T < T_c$
- the most probable configuration for different nucleon number ${oldsymbol{A}}$

• the free energy

• chemical potential $\mu = \partial F / \partial A$



SUMMARY & OUTLOOK

- Nuclear lattice simulations: a new quantum many-body approach
 - \rightarrow based on the successful continuum nuclear chiral EFT
 - \rightarrow a number of highly visible results already obtained
- Algorithmic developments
 - \rightarrow shuttle algorithm speeds up the simulations considerably
 - \rightarrow pinhole algorithm allows to fix the center-of-mass: charge distributions etc
 - \rightarrow pinhole trace algorithm: thermodynamics with fixed nucleon number
- Towards heavier nuclei & higher precision
 - \rightarrow highly improved LO action based on SU(4) \rightarrow nuclei & neutron matter
 - \rightarrow NN interaction at N3LO, first results for nuclei at N3LO
- Ab initio nuclear thermodynamics
 - \rightarrow partition function via the pinhole trace algorithm
 - \rightarrow first promising results for the phase diagram of nuclear matter at finite temperature
 - \rightarrow prediction of the vapor-liquid phase transition within reasonable accuracy